

ESS 415/515

Erika Harnett:

Office: Johnson 267

Phone: 543-0212

Email: eharnett@u.washington.edu

Website: www.ess.washington.edu/Space/ESS515

Grading:

About 5 homework assignments (50%)

Two exams (20%): November 4th and December 9th

Paper/presentation (10%)

Book:

Basic Space Plasma Physics

By Baumjohann and Treumann

Topics Covered:

Introduction to Magnetospheres

Characterizing Plasmas

Generation of the solar wind

Magnetohydrodynamics (and extension to other methods)

Diffusion of Magnetic Field in a Plasma

Intro to Waves and Instabilities in a Plasma

Diamagnetism and Boundaries

Guiding Center Theory

Magnetic Mirroring and the Radiation Belts

Gradient and Curvature Drifts

Currents in the Magnetosphere

Substorms, storms and aurora

Plasma processes at other planets and moons

Related topics (dynamo theory, lab plasmas)

Paper

Due December 2, 2009

Write a 5-6 page paper on the important processes in the near space environment of the planet/moon of your choice other than the Earth. This means the exosphere and above for of a planet /moon with a thin atmosphere, or the ionosphere and above of a planet/moon with a thick atmosphere. Discuss how these processes combine with the inherent characteristics of the planet/moon, such as the composition, distance from the Sun/star, or location within a planet's magnetosphere, to make the system unique and/or of scientific interest.

Use 1.5 spacing, 10-12 pt font, and one inch margins. Figures are allowed and encouraged but they can not constitute more than 1 full page of the paper and make sure you cite the source of the figure.

You must use at least three refereed journal articles as sources of information (e.g. from Journal of Geophysical Research, Astrophysical Journal, Science). You can use (reputable) web sites for additional information; just make sure you cite them like a regular reference. Please list your reference in your bibliography according to the following style:

Author #1, Author #2, “Paper Title”, *Journal Name*, vol., (no.), pages, year

Some web sites to help you get started are:

<http://solarsystem.nasa.gov/planets/index.cfm>

<http://www.nineplanets.org>

http://www.jpl.nasa.gov/solar_system/planets/planets_index.html

Your paper will be graded on

- scientific accuracy,
- the clarity of your explanations/discussion, and
- grammar.

You will also give a 12 minute presentation in class, on December 2nd or 7th, on the material.

Sun

Solar wind \Rightarrow boiling off stellar atmosphere
(corona \sim 1 million degrees)

H^+ & e^- \sim 95%

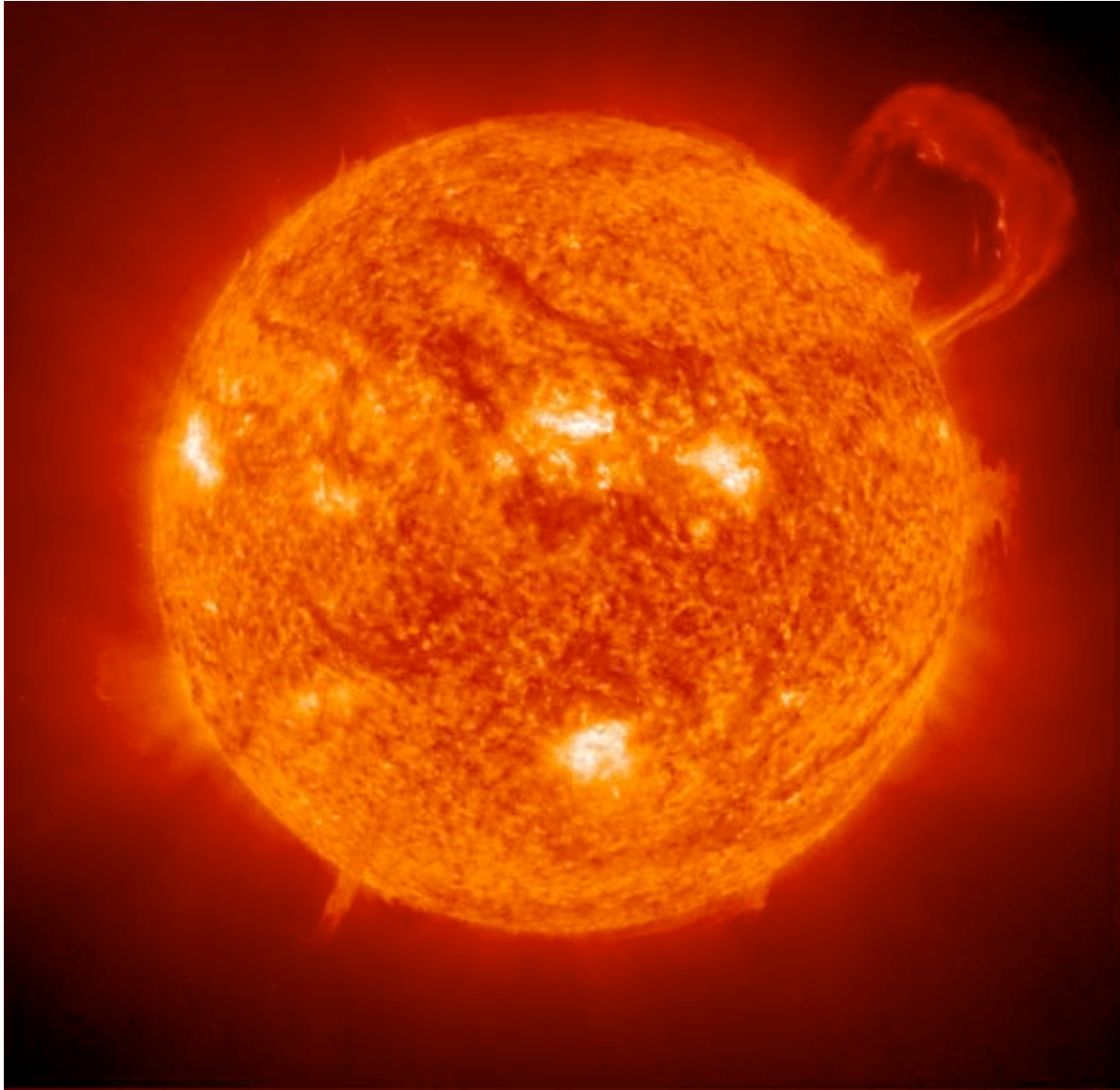
He^+ \sim 4 – 5%

O^{6+} rare but a tracer in magnetosphere

n \sim 5/cm³

V \sim 400 km/s \Rightarrow Sound speed \sim 50 km/s





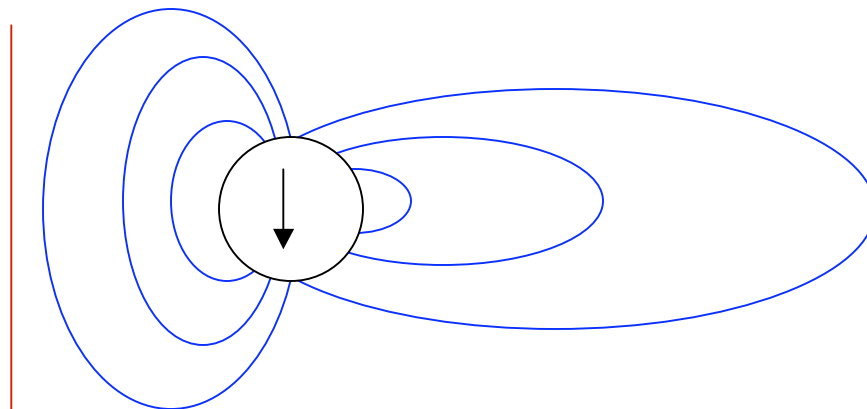
Magnetic Field

Active Regions

- Coronal Mass Ejections
 - High Density
 - Turbulent Magnetic Field
- High speed flows



SW + magnetic field \Rightarrow magnetosphere

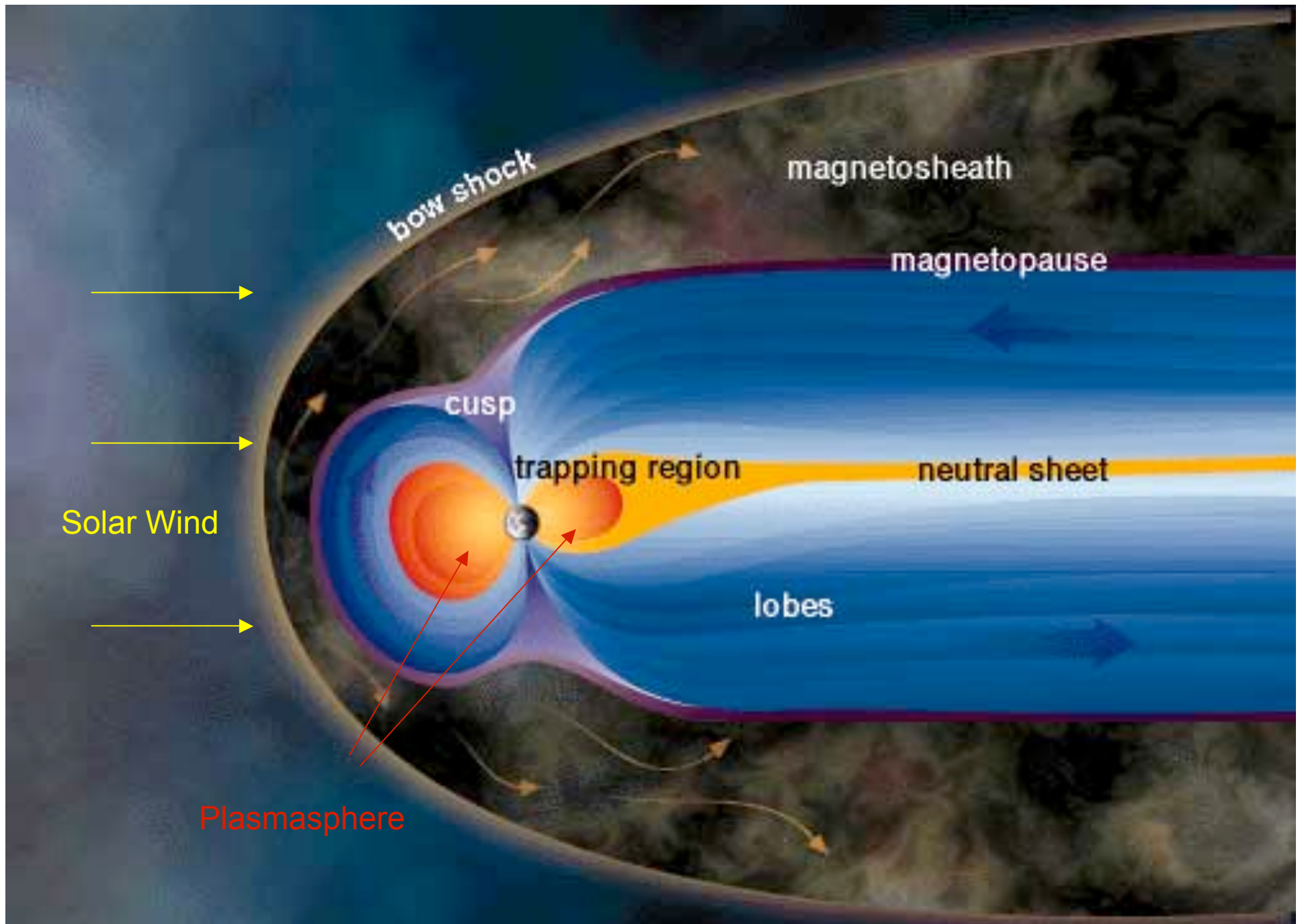


Bow Shock : solar wind supersonic

Magnetopause : balance region; SW dynamic pressure = planet magnetic pressure

$$\rho v^2 = \frac{B^2}{\mu_0}$$

Plasmasphere : under Earth's influence (continuation of ionosphere)



Planetary Magnetic Fields

Earth, Jupiter, Saturn, Neptune,
Uranus, Mercury (small),
Ganymede

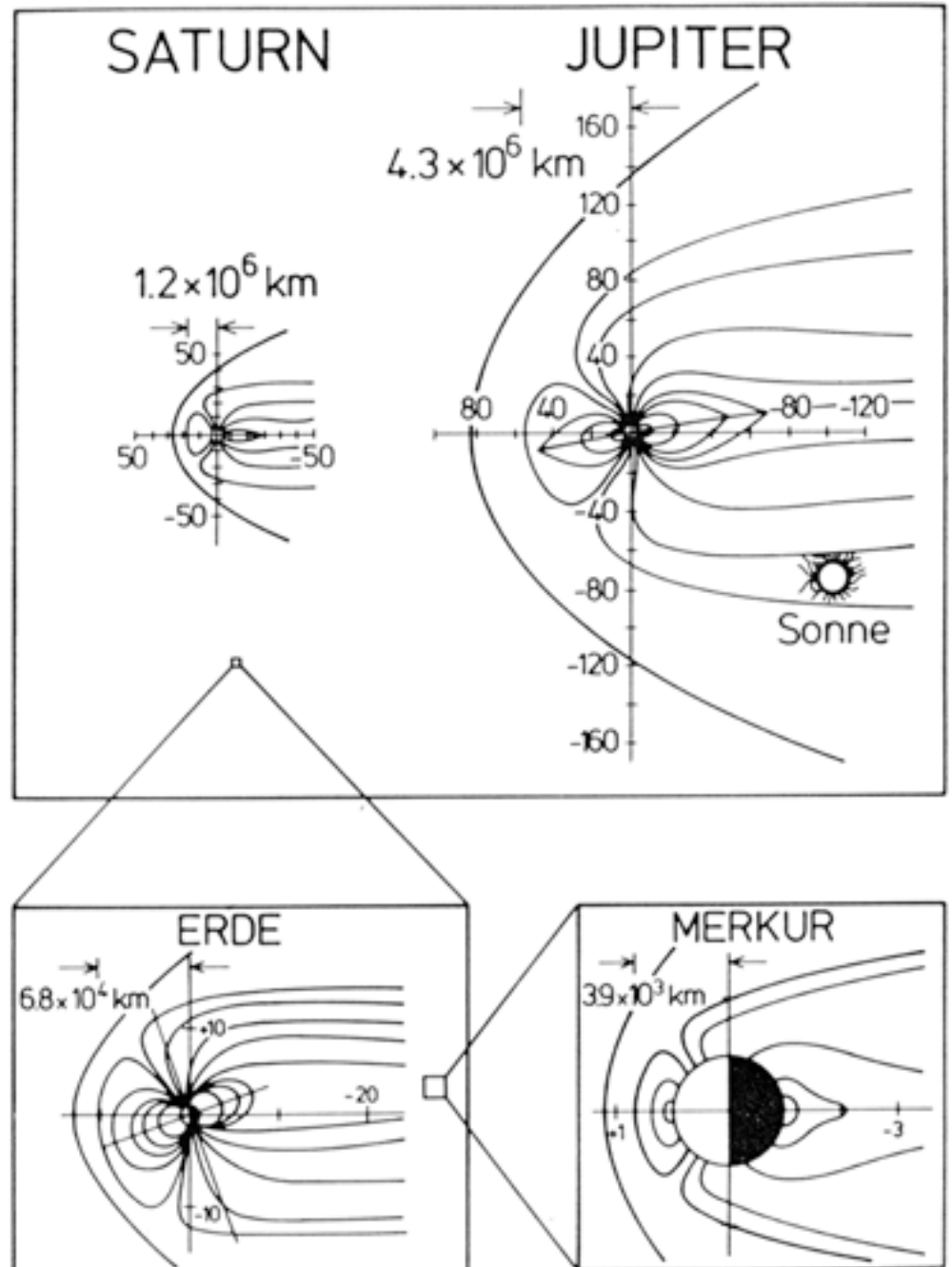
No Global Field

⇒ induced magnetic fields

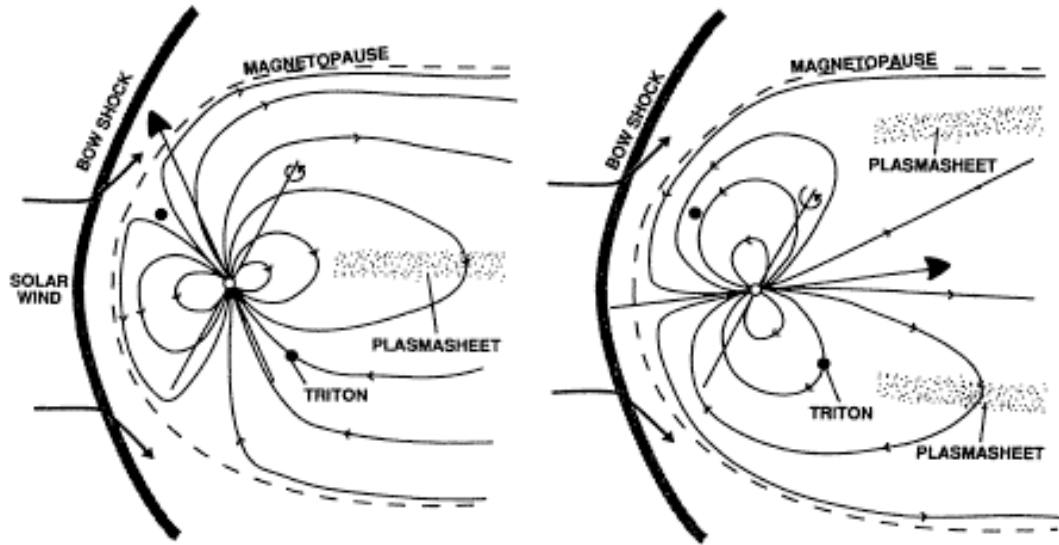
Venus, Io, Europa

Magnetic Anomalies

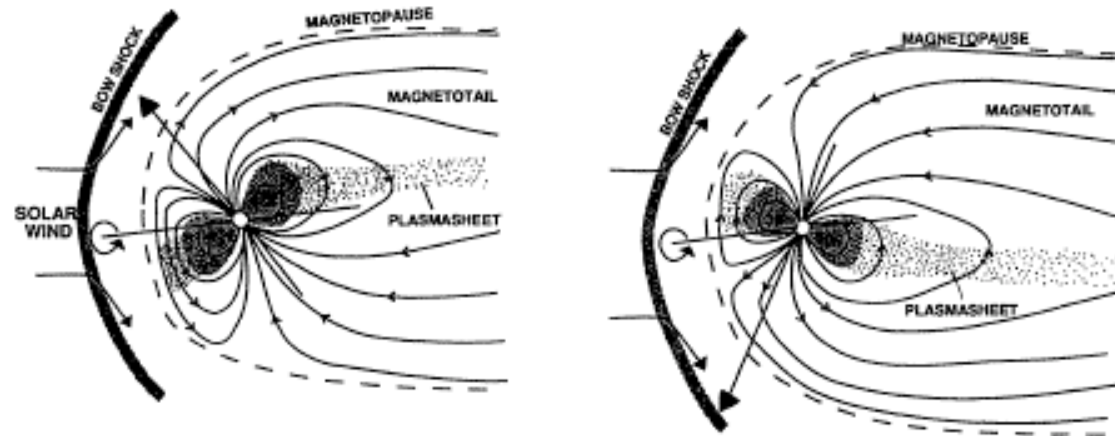
Mars, Moon



NEPTUNE



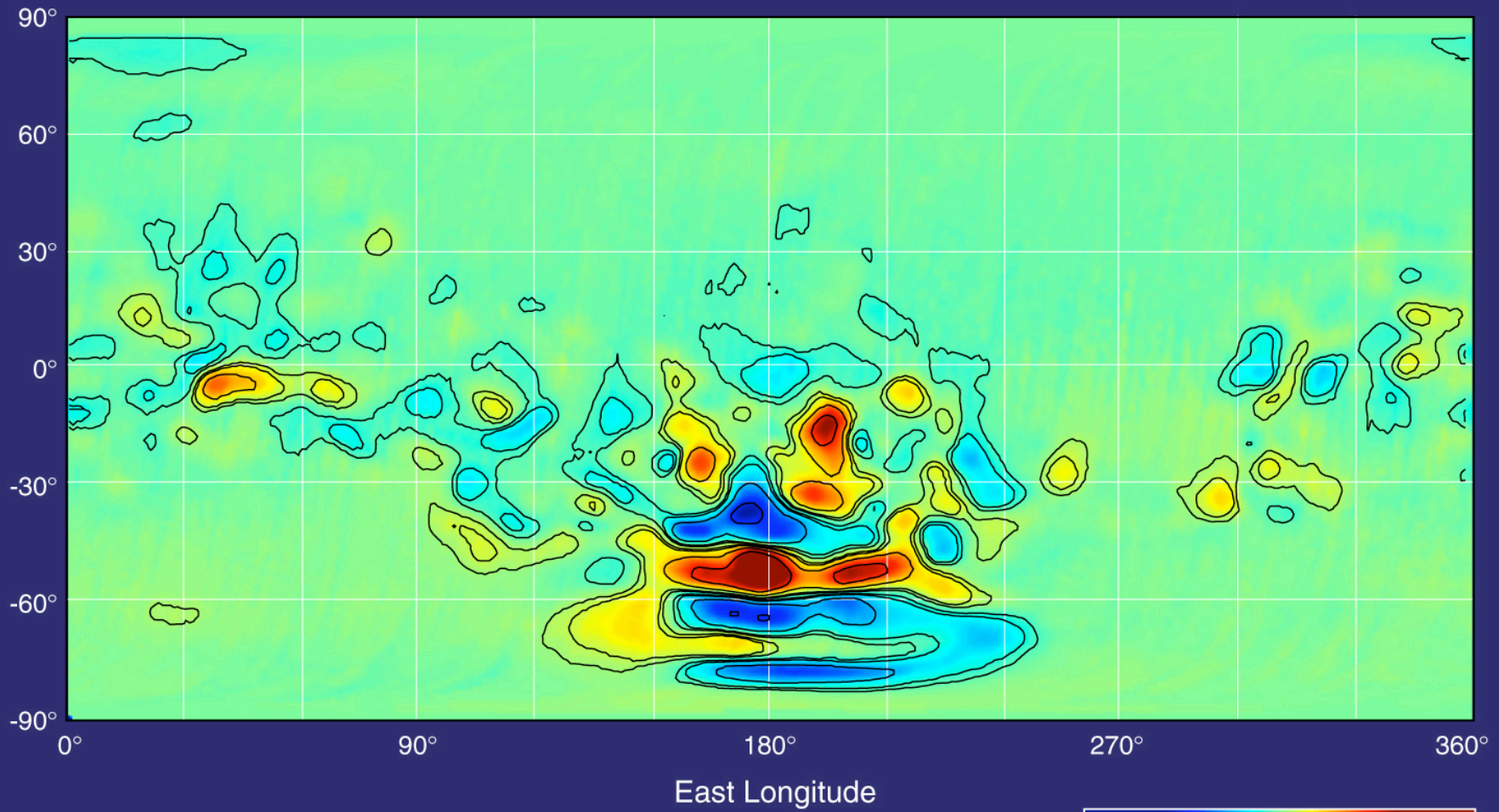
URANUS



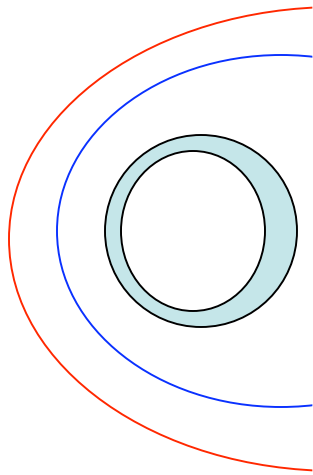
Mars Crustal Magnetism

Mars Global Surveyor

MAG/ER



SW + ionosphere \Rightarrow



Ionopause = balance region (thermal pressure of ionosphere)
 $\rho v^2 = nkT$

MPB = magnetic pileup boundary

Bow shock

Applicable to:

Mars, comets, Pluto(?)

Ion pick-up – applicable to Jovian moons

Why is this important?

- Space Weather
- Planetary Evolution
- Overlap with other fields